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LETTER TO THE EDITOR

Coarse graining, Monte Carlo renormalisation, percolation threshold and critical temperature in the Ising model

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Abstract. Finite-size scaling suggests that the super-spins corresponding to large cells in a Monte Carlo renormalisation scheme form at $T = T_c$ an infinite network of connected down spins. Our finite-size scaling assumption, $M_b = f[(T - T_c)b^{1/\nu}]$ for the renormalised block spin magnetisation is tested by computer simulation for lattice sizes 128^3 and 256^3 .

Clusters in Ising models or lattice gases are usually defined as groups of down spins connected by exchange forces J . If one approaches the Curie temperature T_c of the Ising magnet from below, and looks at the phase where most spins point up, then in general the percolation temperature T_{perc} , where the down spins form for the first time an infinite connected network, does not coincide with T_c . For example, in the nearest-neighbour simple cubic Ising model, T_{perc}/T_c is about 0.96, with a magnetisation M^{perc} of about 0.6 at this percolation threshold (Müller-Krumbhaar 1974, Heermann and Stauffer 1981, Kertész *et al* 1983). Only in the modified cluster definition of Coniglio and Klein (1980), where bonds between up spins are formed randomly with probability $1 - \exp(-2J/k_b T)$, is an infinite network of up spins connected by these additional bonds formed at the correct T_c where the magnetisation vanishes.

No such Coniglio-Klein bonds are taken into account in usual coarse graining theories (see for example Bruce and Wallace 1983) which work with a local magnetisation thought to come from averages over cells of linear dimension b . This averaging is done explicitly in some Monte Carlo renormalisation methods (Swendsen 1982, Pawley *et al* 1984, Jan *et al* 1983, Kalle 1984) where the spins in a cell of length b are replaced by a single superspin ± 1 having the orientation of the majority of the primary spins in that cell (with random tie breaking). Our question now is: Do these superspins percolate at the correct T_c , i.e. is an infinite network of down cells formed for the first time at $T_{\text{perc}} = T_c$? We will argue that for finite cells this is not the case but that the positive difference $T_c - T_{\text{perc}}(b)$ between the critical and the percolation temperature vanishes as $b^{-1/\nu}$ for $b \rightarrow \infty$.

We look at the renormalised magnetisation M_b of cells with b^d spins in a d -dimensional (hypercubic) lattice, $d < 4$, of linear dimension L , with $1 \ll b \ll L$. For $b = 1$, M_b simply gives the usual magnetisation of the primary spins. Our finite-size scaling assumption is

$$M_b = f[(T - T_c)b^{1/\nu}] \quad (T \rightarrow T_c, b \rightarrow \infty) \quad (1)$$

where the correlation length $\xi \propto (T_c - T)^{-\nu}$ is assumed to be much smaller than L , but can be smaller or larger than b . We do not assume instead

$$M_b = b^{-\beta/\nu} f[(T - T_c)b^{1/\nu}] \quad (2)$$

where β is the critical exponent for the spontaneous magnetisation. Equation (2) would be correct (Fisher 1971, Landau 1976, Binder 1981) for the usual magnetisation measured in a system of size b ; then $b = \infty$ is the desired thermodynamic limit.

In our case, however, $b = 1$ and not $b = \infty$ corresponds to the usual magnetisation, whereas $M_b \rightarrow 1$ for $b \rightarrow \infty$ at fixed T below T_c . For in a large enough cell below T_c , $\xi \ll b$, the relative magnetisation fluctuations are negligible, the majority of spins point upwards, and thus $M_b = 1$. Figure 1 shows qualitatively the difference between equation (1) for our renormalised magnetisation M_b and equation (2) for the usual magnetisation in a system of finite size b ; if $b = 1$, in the latter case the 'magnetisation' of the single-spin system is always ± 1 . Our renormalised magnetisation M_b of equation (1) therefore is not analogous to the magnetisation in small systems, equation (2). Instead it corresponds to the probability R of percolative systems (Reynolds *et al* 1980, Stauffer 1985) to have a cluster connecting top and bottom in a cell of size b ; this probability R also approaches zero or unity for large enough cells and follows the analogue of equation (1).

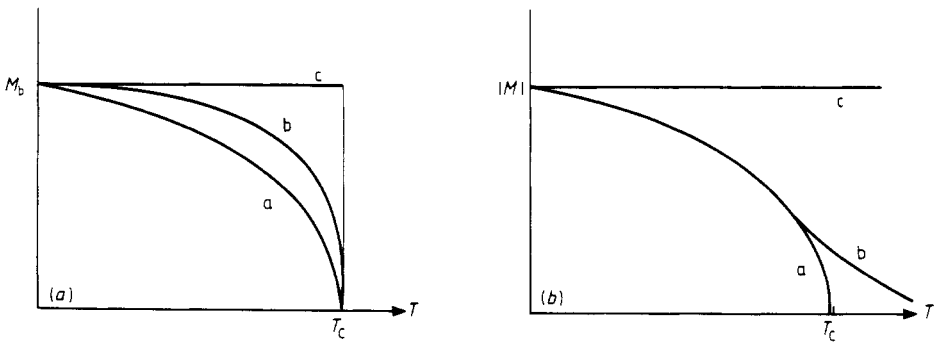


Figure 1. Qualitative comparison of renormalised magnetisation due to cells of size b in a much larger system of size L (a) equation (1) ((a, $b = 1$; b, b large; c, $b = \infty$) and of the usual magnetisation in a finite system of size L (b) equation (2)) (a, $L = \infty$; b, L large; c, $L = 1$).

For the unrenormalised primary spins ($b = 1$) the percolation threshold is reached at $T = T_{\text{perc}}(b = 1)$ below T_c , with a positive magnetisation $M = M_{b=1}^{\text{perc}}$; for the clusters formed by upward oriented neighbouring b cells the percolation threshold is at $T_{\text{perc}}(b)$ and M_b^{perc} . Experience with dynamic renormalisation (Jan *et al* 1983, Kalle 1984) warns against assuming $M_{b=1}^{\text{perc}} = M_b^{\text{perc}}$ and suggest instead, similar to percolation (Reynolds *et al* 1980):

$$M_b^{\text{perc}} = M_{b'}^{\text{perc}} \quad (b \text{ and } b' \gg 1). \tag{3}$$

Equations (1) and (3) give

$$[T_c - T_{\text{perc}}(b)]/[T_c - T_{\text{perc}}(b')] = (b/b')^{-1/\nu} \tag{4}$$

for large enough cells, and thus

$$T_c - T_{\text{perc}}(b) \propto b^{-1/\nu} \quad (b \rightarrow \infty). \tag{5}$$

The reason behind equations (3)–(5) is the similarity of block spins and primary spins on which scaling and real space renormalisation are based: The correlations

between primary spins and between cells are about the same apart from factors of order unity, provided the distance is measured in terms of the correlation length ξ . Therefore the concentration M^{perc} of up cells at the percolation threshold T_{perc} is the same, whether the cells are large or very large.

Direct Monte Carlo evaluation of $T_{\text{perc}}(b)$ might be difficult for large b and should be combined (Binder, private communication) with a detailed comparison of cluster size distributions for the renormalised superspins and the Coniglio-Klein clusters of primary spins. As a first step, we test here the underlying scaling assumption (1) on a CDC Cyber 205 vector computer. To satisfy $1 \ll b \ll L$ our system has to be large. We thus made up to 10^5 Monte Carlo steps per spin for $L = 128$ in a simple cubic lattice between $0.82T_c$ and $0.999T_c$, as well as shorter runs for $L = 256$, using Kalle's program. Our equilibrium results (at $0.999T_c$ extrapolated to $L = \infty$) are shown in figure 2. There $b = 1$ slightly deviates from the scaling assumption (1), as expected, but larger b confirm the similarity rule (1).

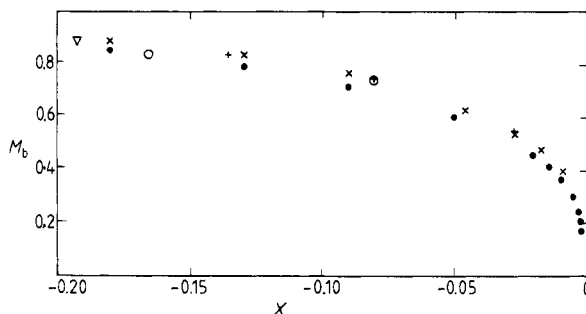


Figure 2. Variation of the renormalised magnetisations M_b with $X = b^{1/\nu}(T - T_c)/T_c$. Our Monte Carlo data (128^3 and 256^3 simple cubic lattices) for $b = 4, 8, \dots$ as listed in the figure follow roughly the same curve and thus confirm equation (1). The factor $b^{1/\nu}$ by which equations (1) and (2) differ varies by roughly a factor 2 between $b = 4$ and $b = 16$; thus our data contradict equation (2). (\bullet , $b = 1$; \times , $b = 4$; $+$, $b = 8$; \circ , $b = 16$; ∇ , $b = 32$).

Thus we see little reason to doubt that the percolation threshold for clusters of Ising superspins in real space renormalisation converges to the desired critical temperature T_c if the cell size b for these superspins goes to infinity. Our result justifies droplet models based on coarse-grained order parameters near T_c and makes feasible future studies of cluster size distributions.

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